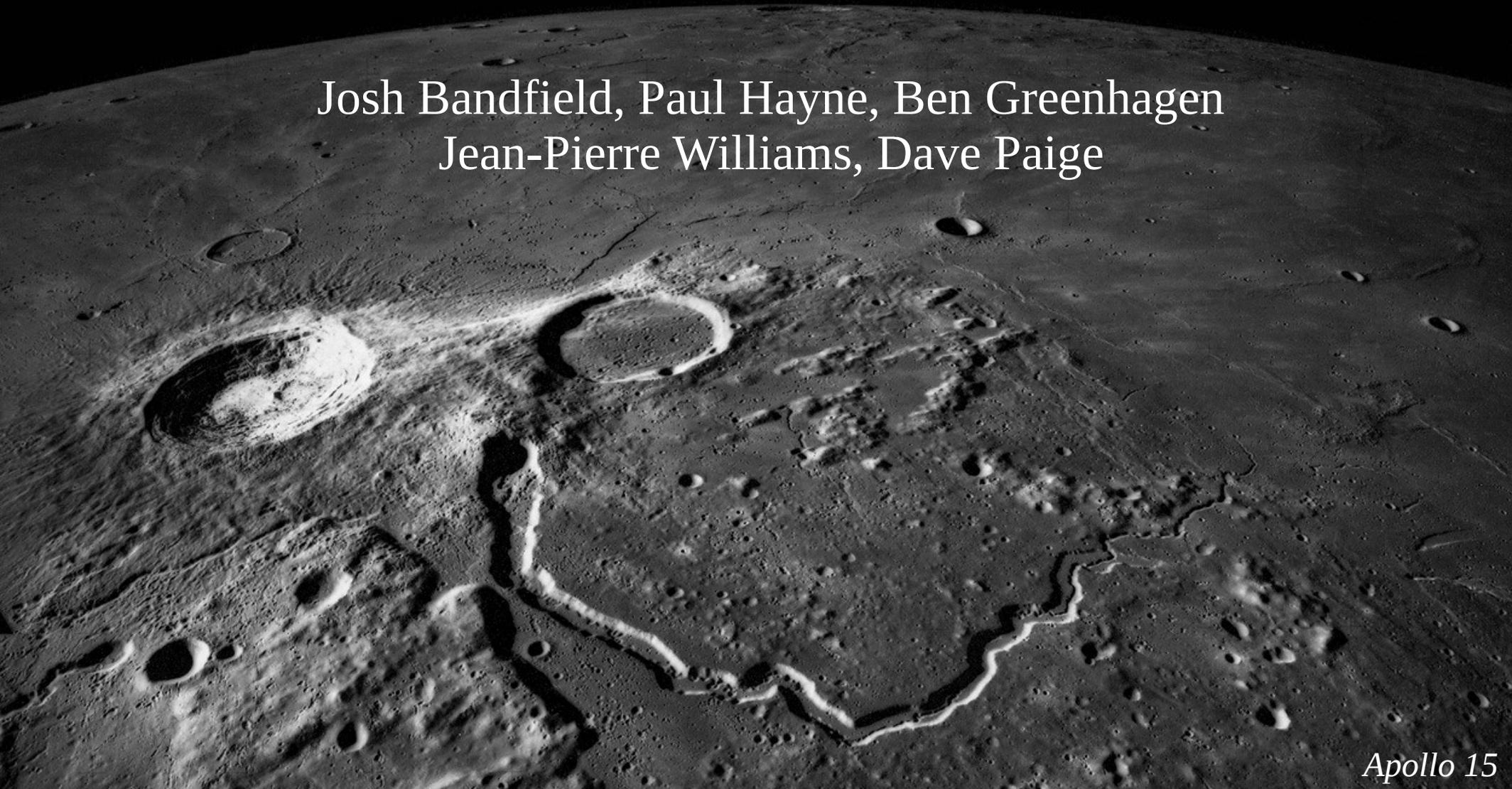


# Getting the temperature right

Thermophysical and spectral properties of planetary surfaces using  
infrared measurements

Josh Bandfield, Paul Hayne, Ben Greenhagen  
Jean-Pierre Williams, Dave Paige



# The Main Point

- Thermal emission from airless bodies is often dominated by effects due to surface roughness and rock abundance
- These surfaces are composed of a wide range of temperatures that cannot be approximated by a single temperature, which can lead to significant errors
- LRO Diviner has begun a campaign to fully characterize thermal emission from the lunar surface, serving as a baseline for investigations throughout the Solar System

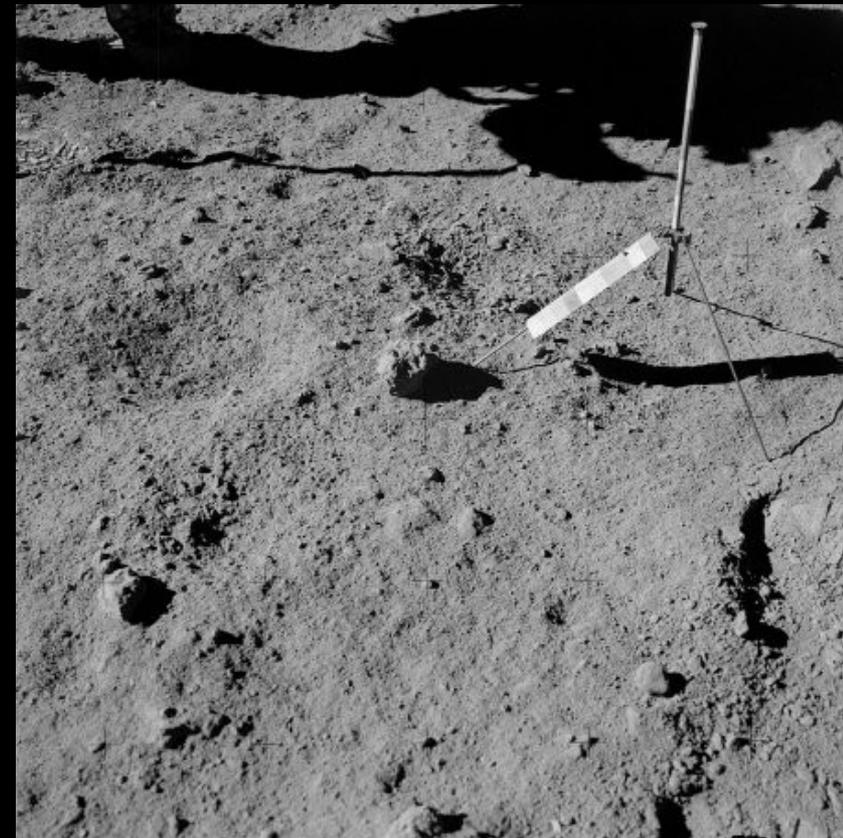
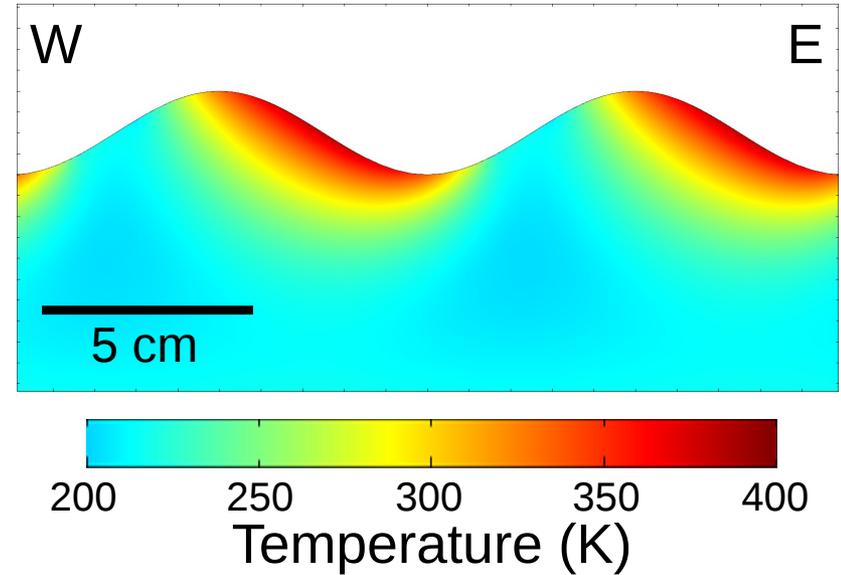
# Surface Roughness

- Small-scale features dominate surface roughness measurement sensitivities
- Surfaces separated by just a few cm can vary by 200 K!

Why??

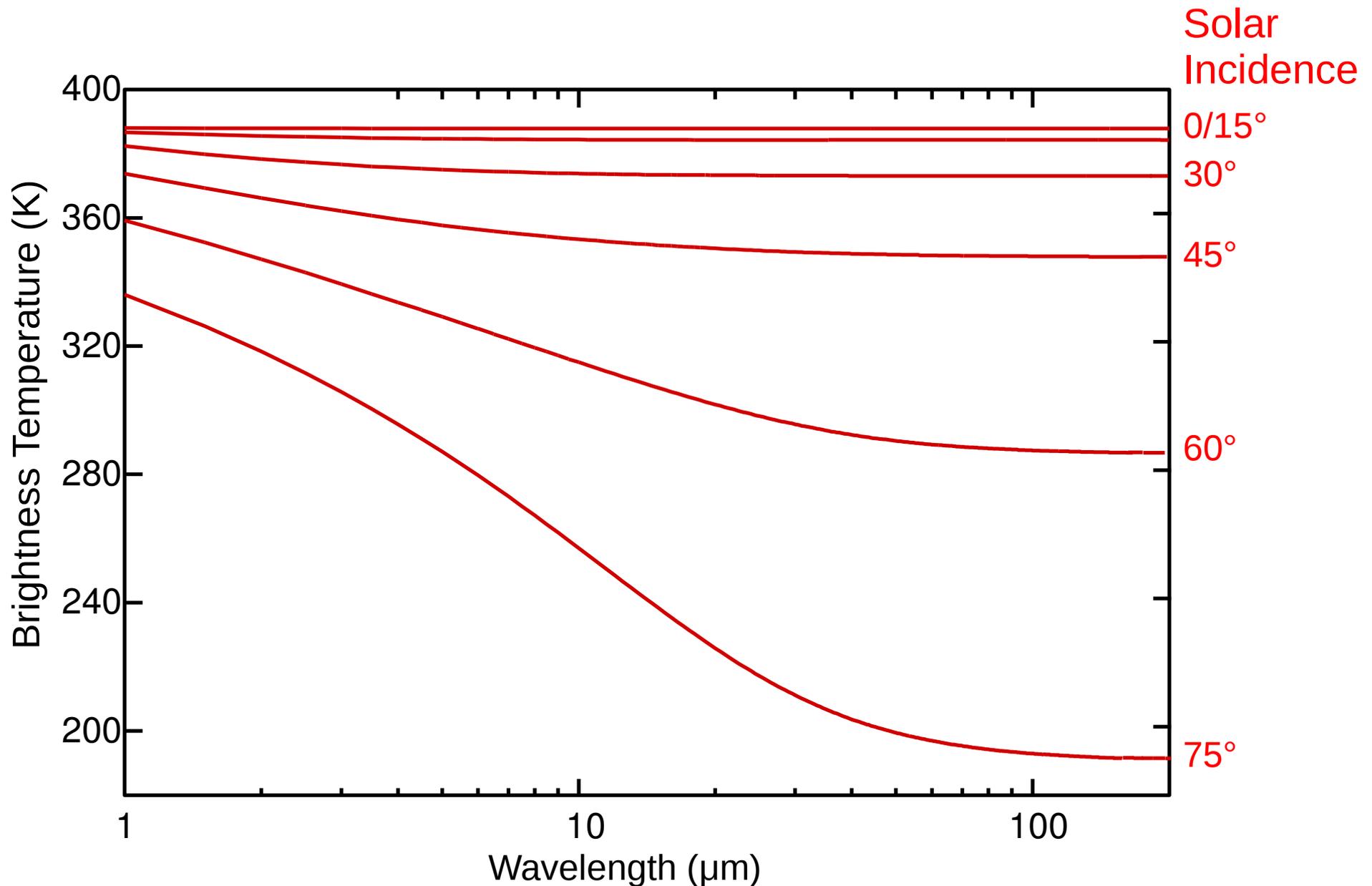
- Rough
- Highly insulating
- Vacuum environment

Modeled 8am equatorial temperatures



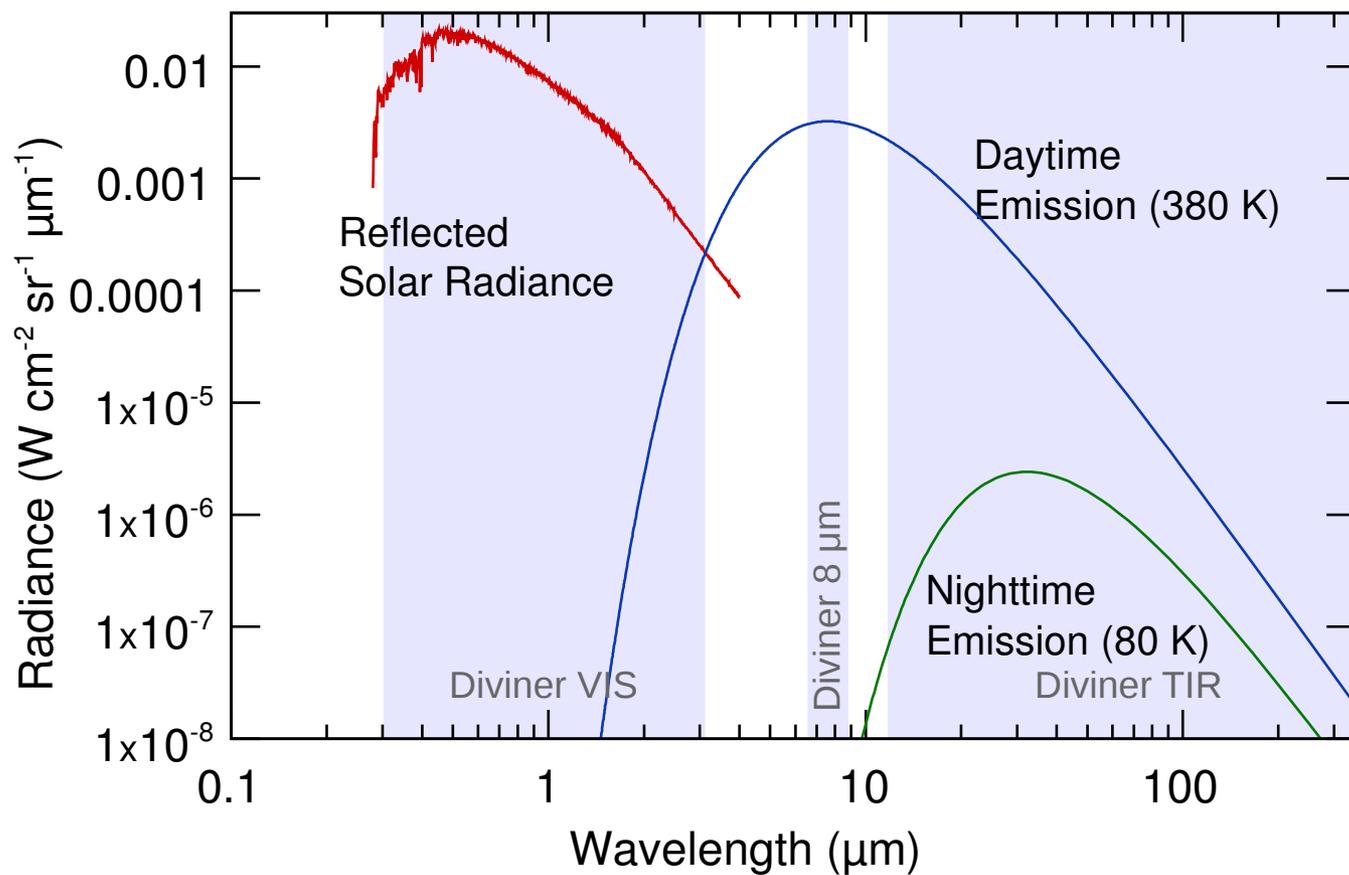
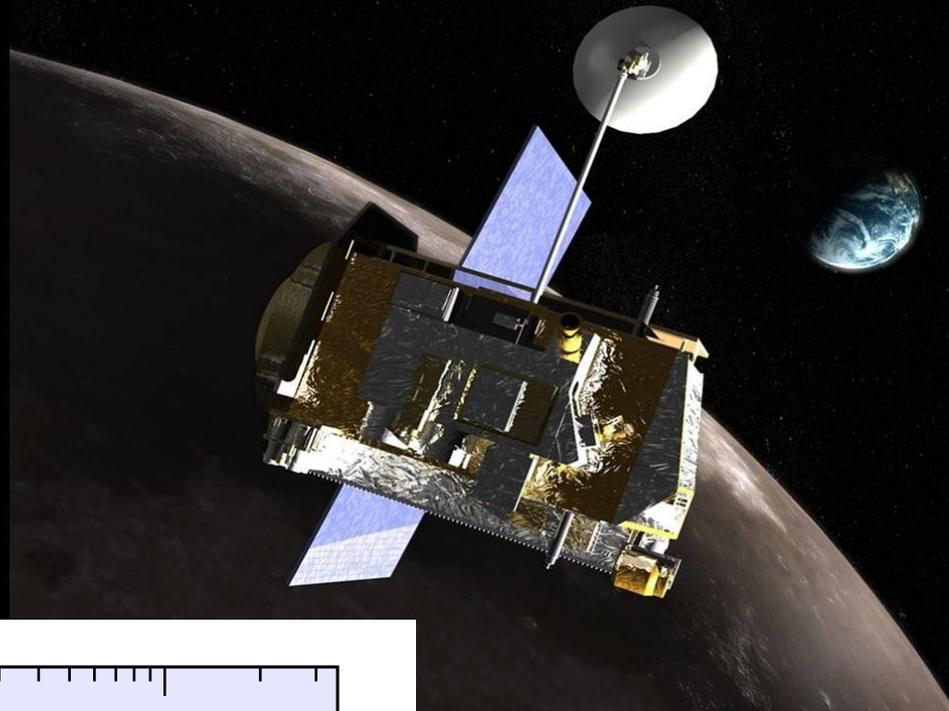
*Apollo 15*

# Surface Roughness and Brightness Temperatures



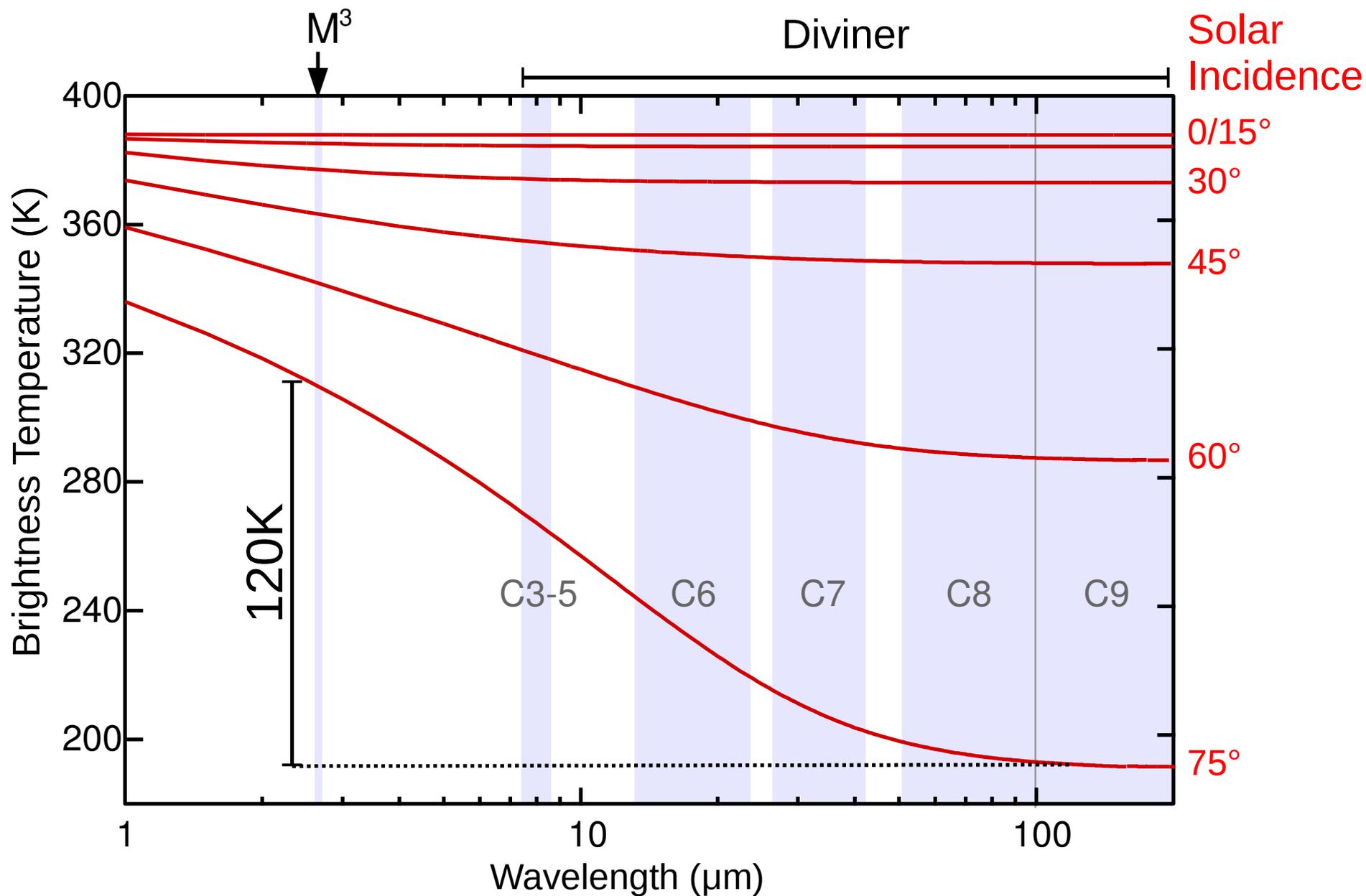
Modeled nadir lunar brightness temperatures for 25° RMS slope distribution

# LRO Diviner: A Uniquely Complete Set of Measurements



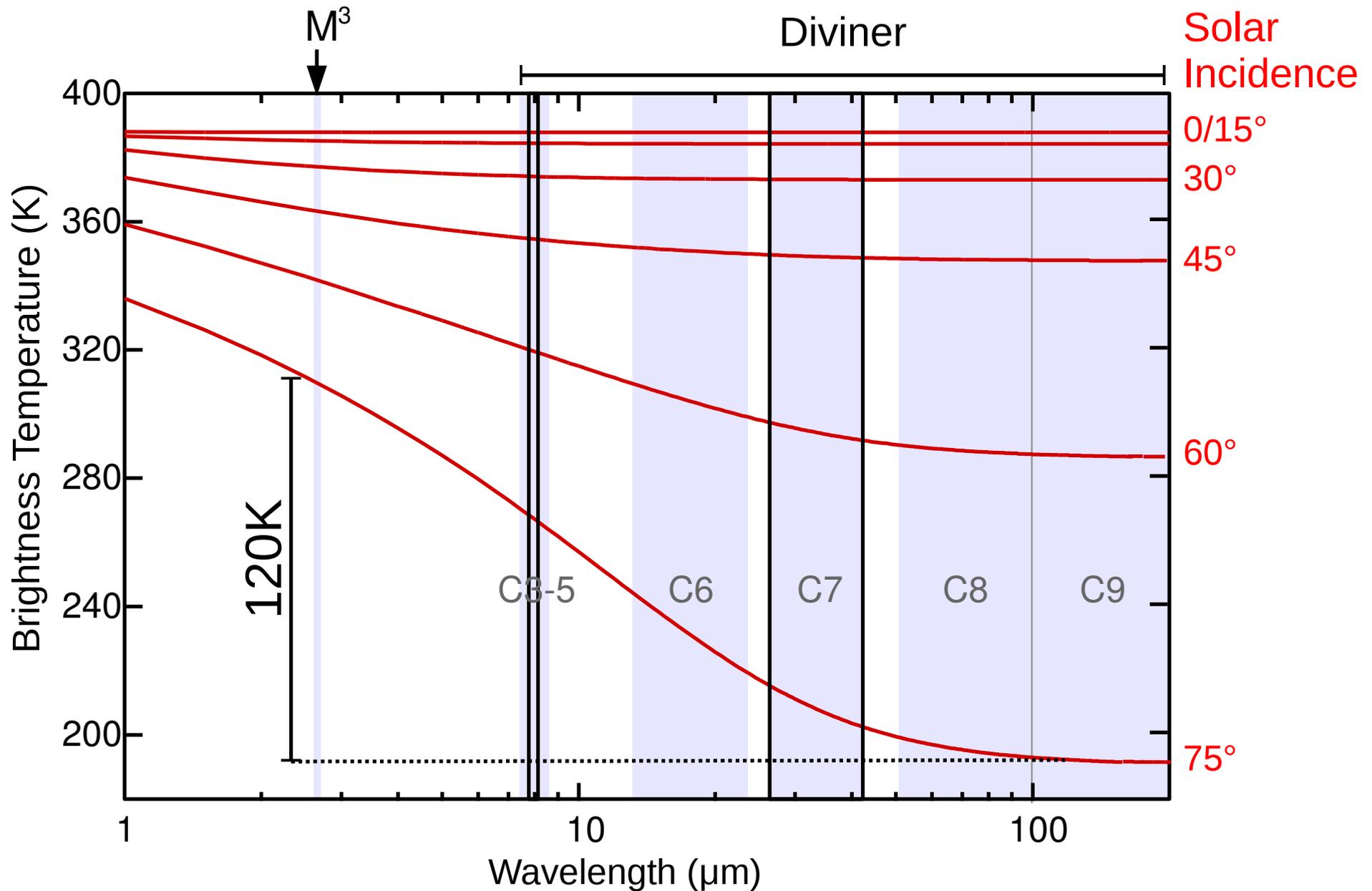
Diviner can measure most emitted and reflected light from most angles at all times of day

# Surface Roughness and Brightness Temperatures



Modeled nadir lunar brightness temperatures for 25° RMS slope distribution

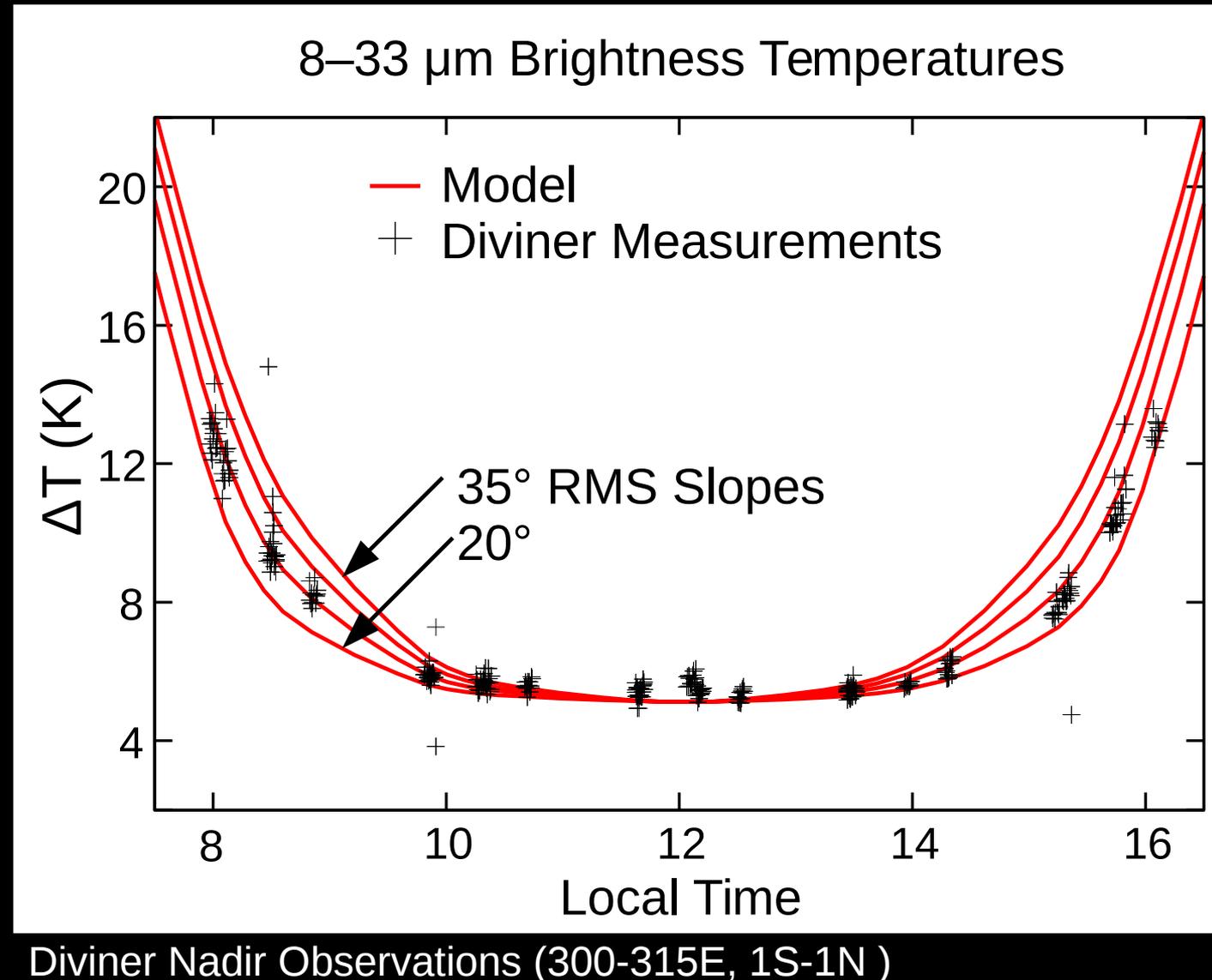
# Surface Roughness and Brightness Temperatures



Modeled nadir lunar brightness temperatures for 25° RMS slope distribution

# Lunar Surface Roughness

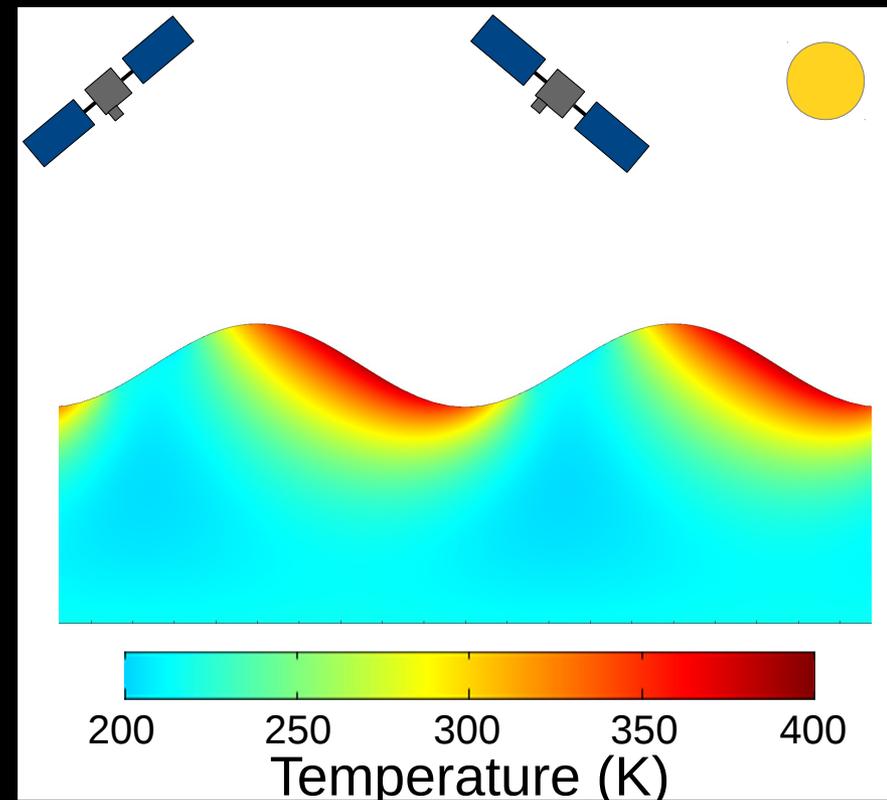
- Anisothermality is closely predicted using a  $\sim 25^\circ$  RMS slope distribution (Bandfield et al., 2015)



# LRO Diviner Phase Function Observations

- Two types of observations:
  - 1) **Nadir spectral observations** – measure spectral differences at a variety of solar incidence angles
  - 2) **Multiple emission angle observations** – measure surfaces from multiple angles at one or many solar incidence angles

- These two observation types provide complementary information and both are being collected by Diviner



# LRO Diviner Observations: Surface Phase Function Measurements

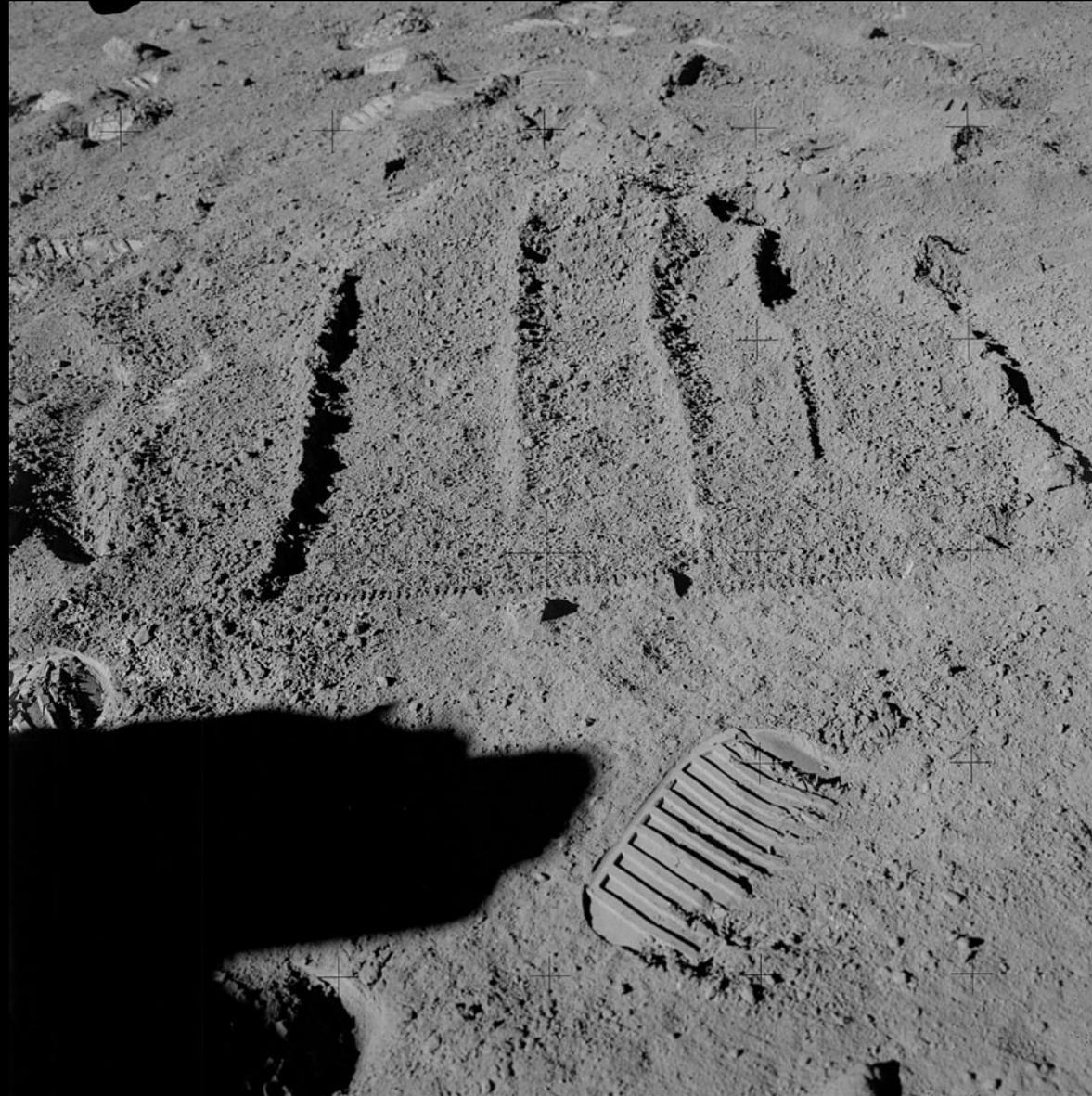
## Goals:

- Define Vis and TIR phase function as completely as possible over representative surface units
  - Full diurnal coverage
  - Up/Down track and previous/post orbit cross-track observations
- Characterize the radiative balance of the lunar surface
  - **TIR coverage will be an unprecedented dataset**
  - There is also significant value in Vis observations that complement LROC
- Map variations in surface properties by collecting sustained off-nadir observations

# LRO Diviner Observations: Surface Phase Function Measurements

## Benefits:

- Determination of cm-scale physical properties to help determine the history and development of the lunar regolith



# LRO Diviner Observations: Surface Phase Function Measurements

## Benefits:

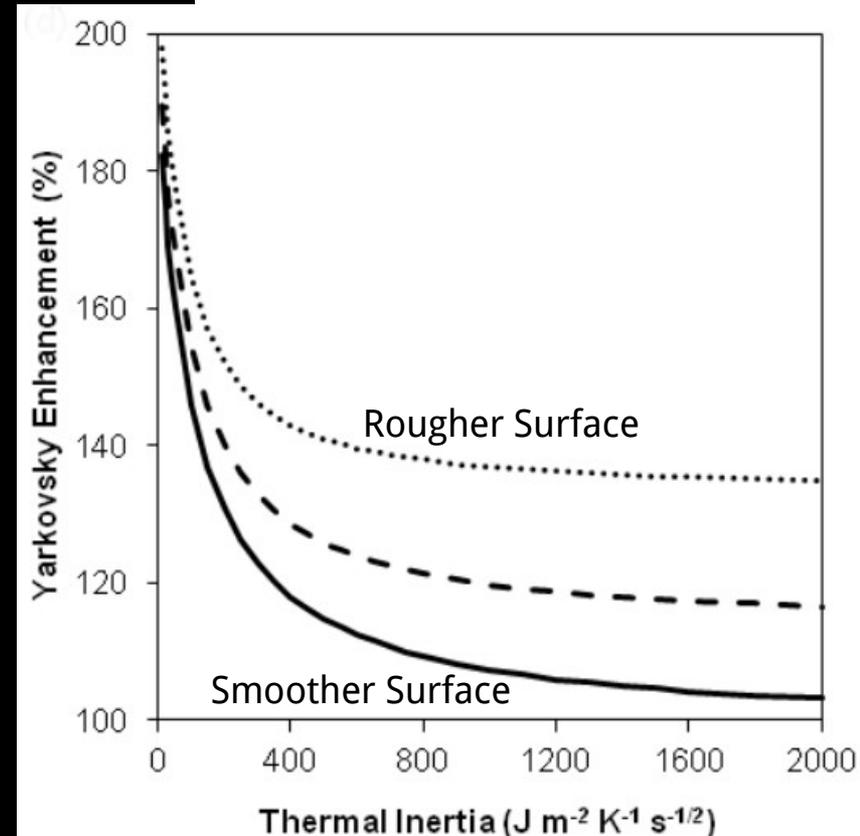
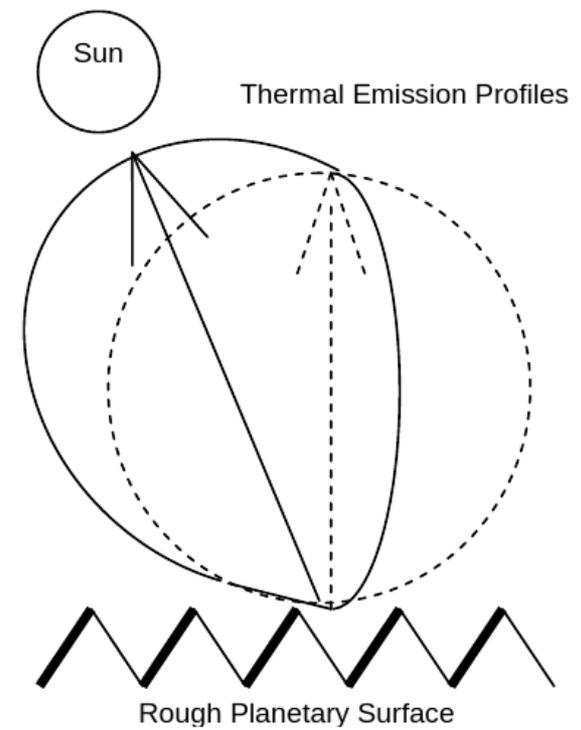
- Improved understanding of disk-integrated and limited measurements of airless bodies
  - Example: Estimation of asteroid size depends on “beaming”, which accounts for roughness properties towards the limb



# LRO Diviner Observations: Surface Phase Function Measurements

## Benefits:

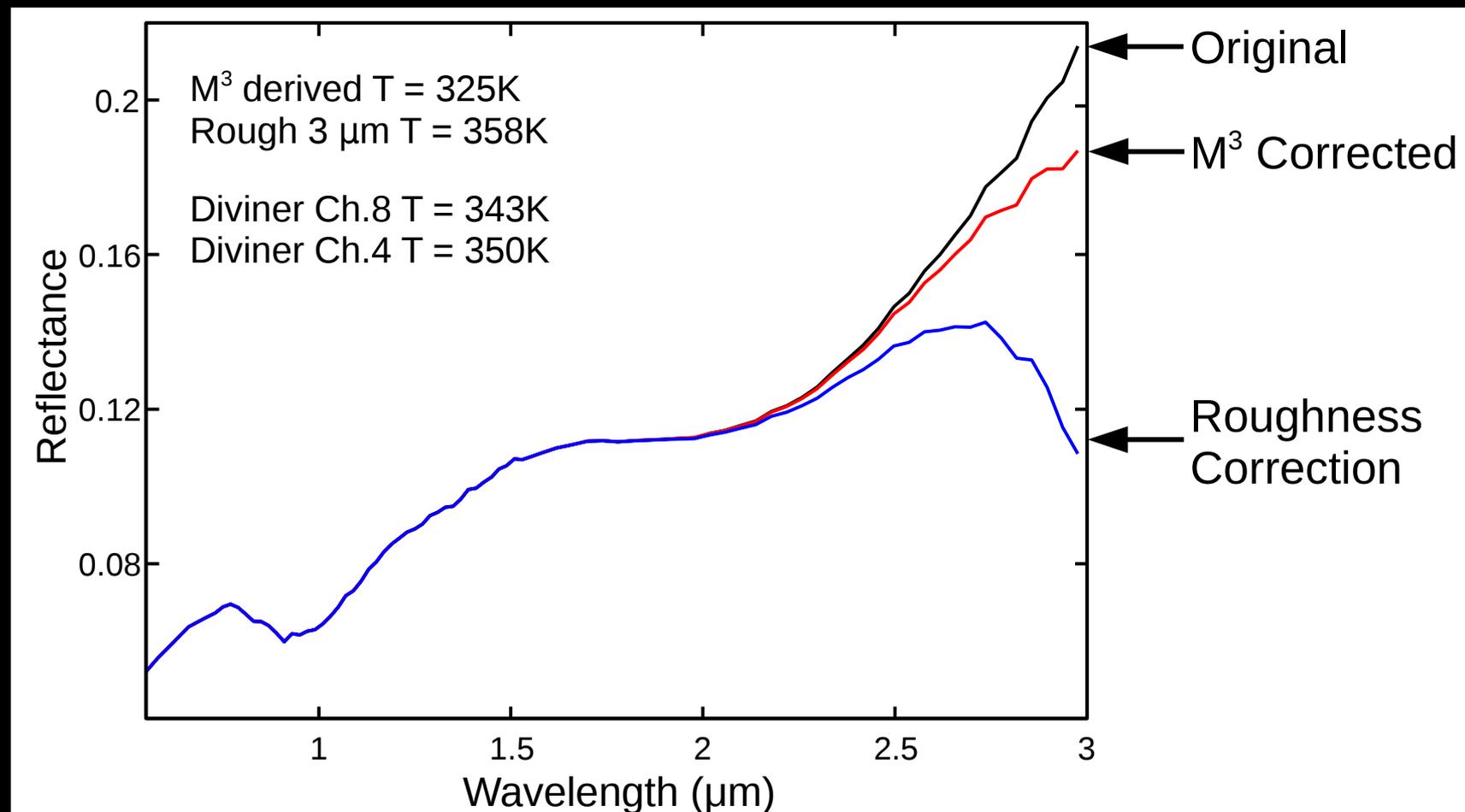
- Improved determinations and understanding of Yarkovsky and YORP effects
  - Example: Non-lambertian emission greatly skews direction of emitted radiance



# LRO Diviner Observations: Surface Phase Function Measurements

**Benefits:** Improved thermal corrections for infrared spectroscopy

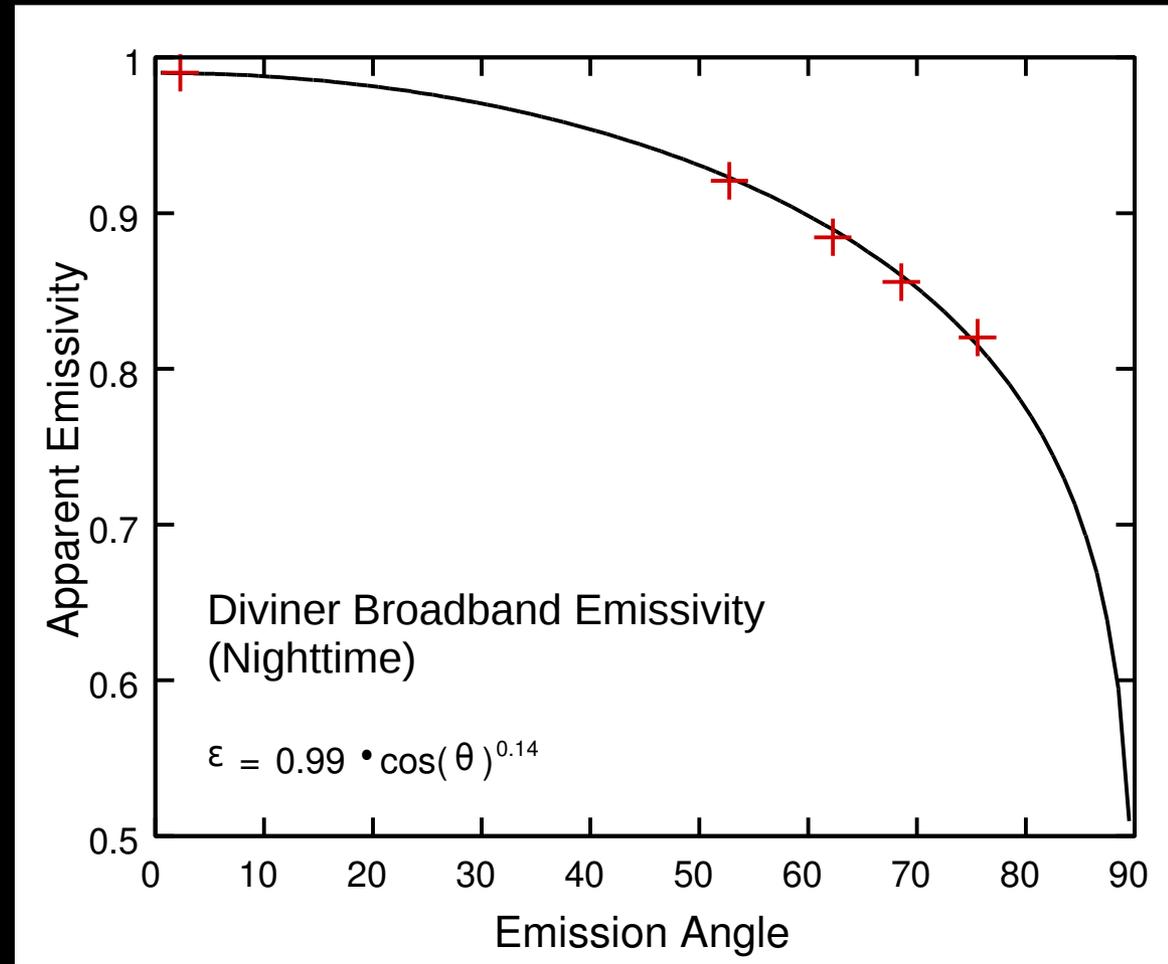
- Example:  $M^3$  OH/ $H_2O$  absorption is much more prominent when accounting for roughness



# LRO Diviner Observations: Surface Phase Function Measurements

## Benefits:

- Characterization of radiative balance for improved thermal modeling
  - Example: Lunar hemispherical emissivity is lower than nadir directional emissivity resulting in warmer temperatures both day and night



# LRO Diviner Observations Can Be Used To Better Understand Many Infrared Datasets



*Dawn at Ceres*

Examples:

## Current/Recent Telescopes

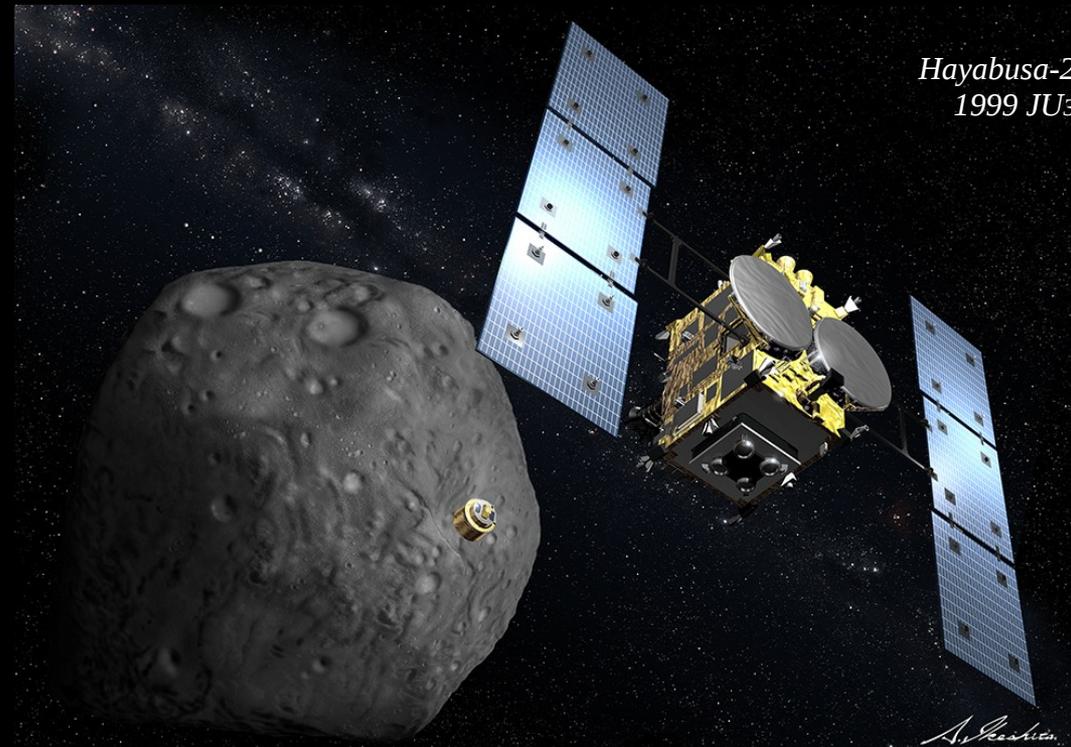
- Spitzer, Herschel, SOFIA, WISE, IRTF

## Current/Recent Spacecraft

- Dawn, LRO, Cassini, MESSENGER, Chandrayaan-1

## Future Spacecraft

- BepiColombo, OSIRIS-REx, Hayabusa-2, Europa Clipper



*Hayabusa-2  
1999 JU3*

*As Perkin*

